



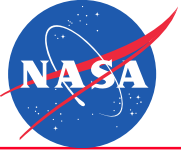
The effect of photon counting detector blocking on centroiding for deep space optical communications

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The work described here was performed at the Jet Propulsion Laboratory, California Institute of Technology, under contract with the National Aeronautics and Space Administration (NASA).

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- **System overview**
 - **Pointing and tracking overview**
 - **Uplink beacon signal**
 - **Background subtraction**
- **Blocking model**
 - **Blocking compensation**
 - **Blocking impact on MSQ statistic**
- **Centroid performance results**
 - **Monte-Carlo simulation**
 - **Laboratory testing**

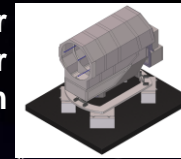
Deep-Space Optical Communications (DSOC)

– OBJECTIVES:

Advance NASA's enhanced communication goals by:

- *Demonstrating optical communications from deep space (> 2 AU) to validate:*
 - *Link acquisition and laser pointing control*
 - *High photon efficiency signaling*

Flight Laser Transceiver (FLT), 22 cm



Psyche Spacecraft (2022)

1064 nm uplink

1550 nm downlink

Ground Laser Transmitter (GLT)
Table Mtn., CA
1m-OCTL Telescope



Ground Laser Receiver (GLR)
Palomar Mtn., CA
5m-dia. Hale Telescope



Deep Space Network (DSN)



Psyche Ops Center



DSOC Ops Ctr.

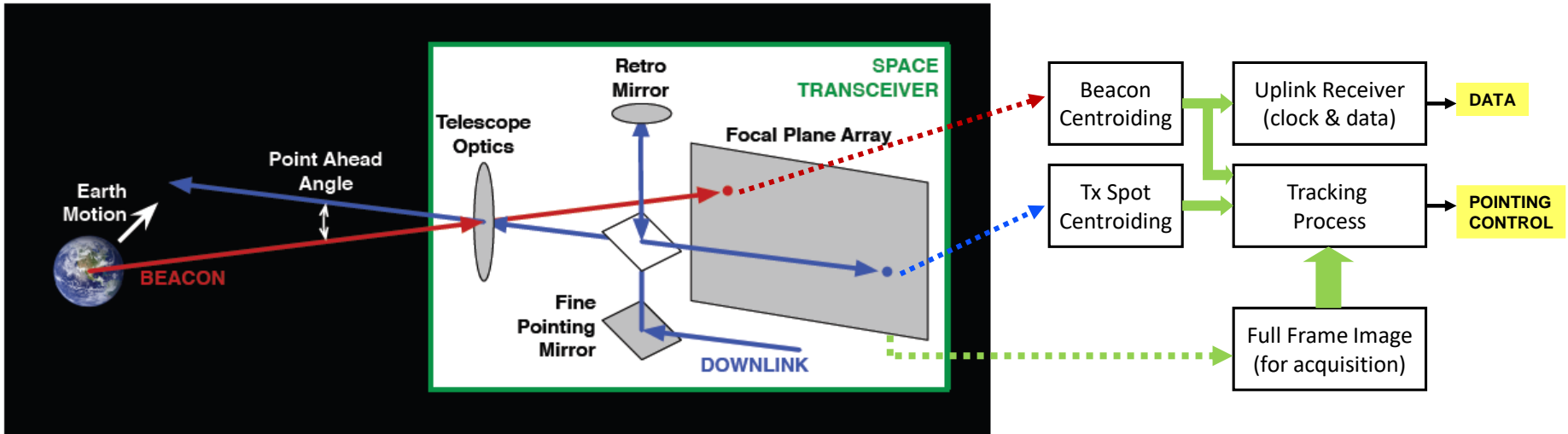


See DSOC presentation by
Abhijit Biswas

Tuesday Jan 30, 11:10-11:30 AM
Room 158

(Paper 10524-30)

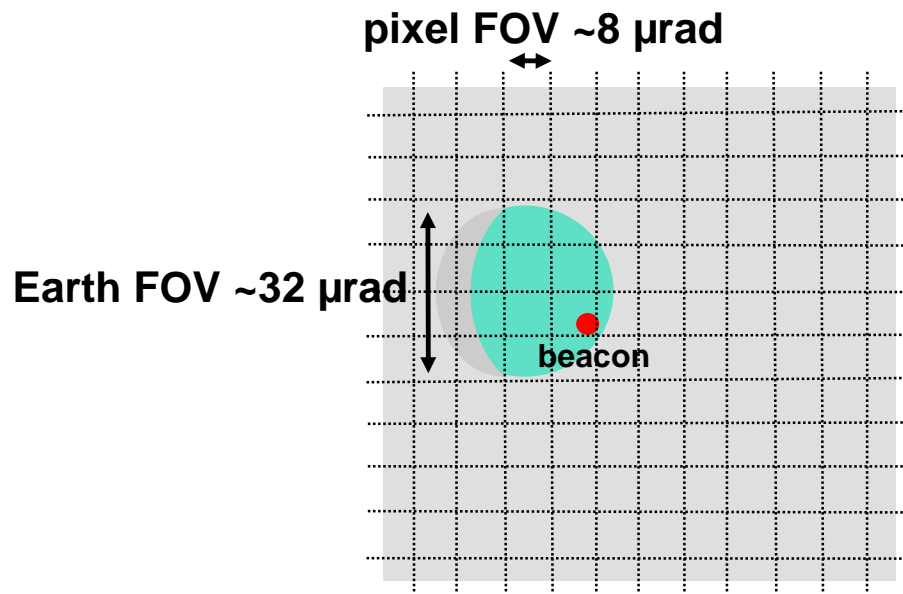
Pre-Decisional Information -- For
Planning and Discussion Purposes Only



- Uplink beacon provides reference pointing information without increasing mass of flight transceiver.
- Uplink signal position is estimated in order to
 - Adjust flight terminal platform attitude and calculate point-ahead angle for downlink transmission
 - Acquire and demodulate uplink data
- A single detector array reduces alignment errors and optical losses
 - Used to estimate location of dim laser beacon to point transmit beam to Earth ground receiver
 - Tracks angle of the transmit beam to confirm the point-ahead angle
 - Photon counting array has best combination of sensitivity and bandwidth
- Signal processing needed for simultaneous spatial acquisition, tracking, parameter estimation and data demodulation

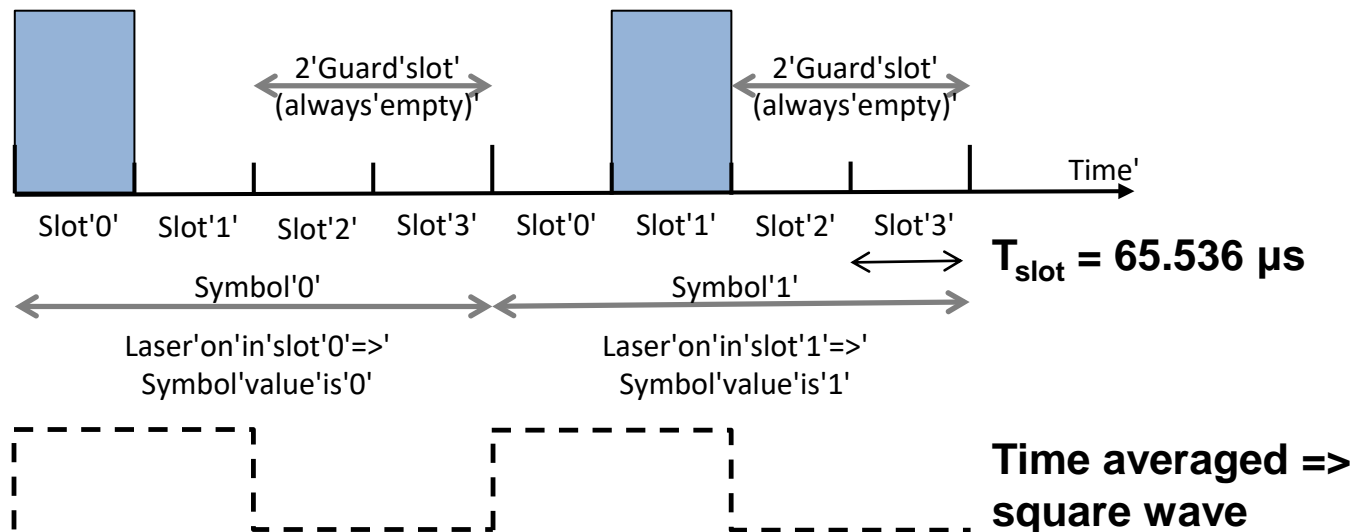


- Uplink beacon tracking error translates directly to downlink pointing error
- $<1 \mu\text{rad}$ beacon centroiding accuracy needed
- Example case: Spacecraft at 2.7 AU



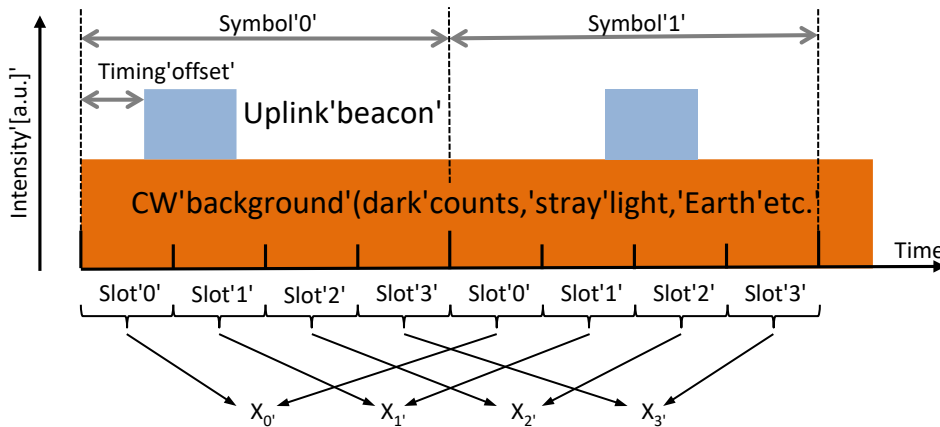
- Beacon irradiance in pW/m^2
- Earth radiance dominates beacon flux
- Beacon centroid biased towards Earth centroid
- We need subpixel beacon centroid accuracy, using currently available photon-counting cameras

- The DSOC uplink beacon uses a 2-PPM with 2 intersymbol guard time (ISGT) slots modulation
- This scheme provides a low rate command channel



- 50% average duty cycle beacon is optimal for signal estimation
- Background subtraction is implemented using a pair of up-down counters offset by command channel slot width

- The modified square law (MSQ) statistic is calculated for each pixel on the array using N symbols ($N=64$ for 60Hz update rate).



$$X_m = \sum_{i=0}^{N-1} x_{corr}(x_{4i+m}), \quad m \in \{0, 1, 2, 3\},$$

$$U = X_0 + X_1 - X_2 - X_3,$$

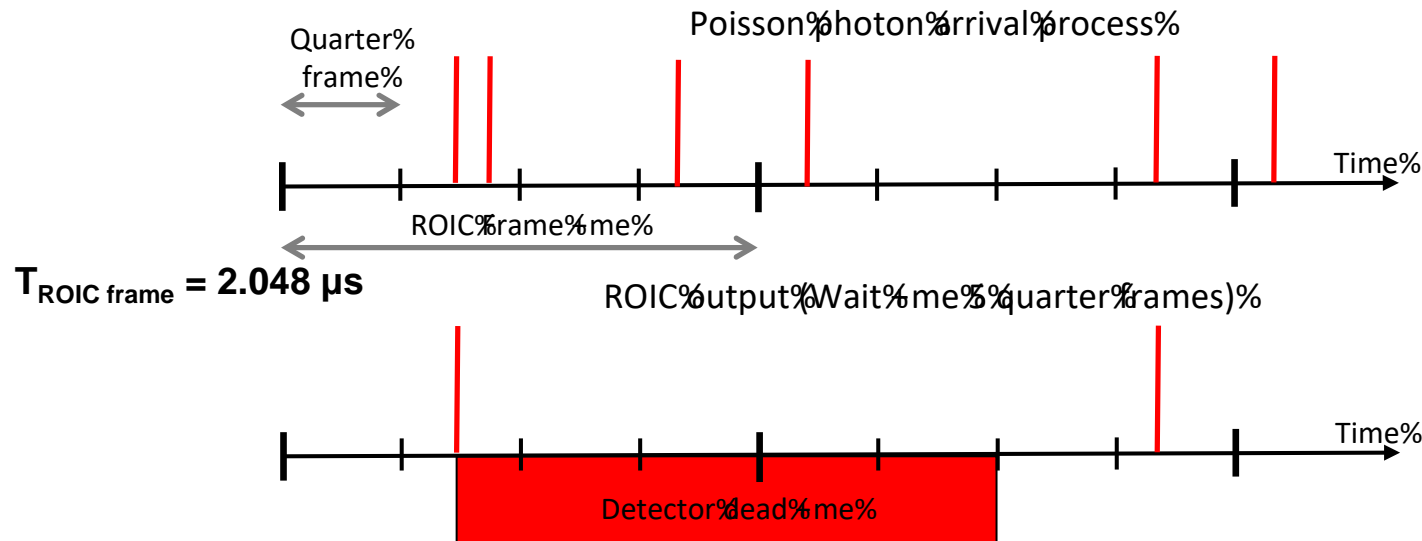
$$V = X_0 - X_1 - X_2 + X_3,$$

$$S = X_0 + X_1 + X_2 + X_3.$$

$$W = U^2 + V^2 - 2S.$$

- The MSQ statistic, W , does not depend on the background level and is proportional to the modulated signal level squared.
- The centroid estimate is calculated using the MSQ statistic in a small subwindow (e.g., 2x2 or 3x3 pixels).

- **Blocking** limits maximum detectable counts, and comprises two effects
 - **Detector dead time** – After a photon detection event the detector is held in an off-state for a set amount of time to limit afterpulsing
 - **ROIC frame rate/single photon detection limit** – at most one count per frame recorded
- The dead/wait time on the PCC ROIC is discretized in units of quarter ROIC frames.
 - The nominal wait time is 5 quarter frames:



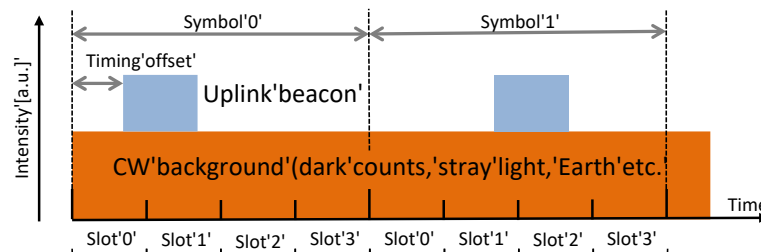
- **Blocking** manifest itself as a nonlinear response at high count rates and is exacerbated by pulsed or modulated signals

- Using straight forward statistics, the effect of blocking can be compensated for.
- The most simple form is: $\lambda_{true} = -\ln(1 - \lambda_{obs})$.

- For the DSOC PCC+ROIC setup, this translates to:

$$\lambda_{true} = -\ln(1 - \lambda_{obs,adj}) = -\ln \left(1 - \frac{x}{n - x(N_{wait} - 1)} \right).$$

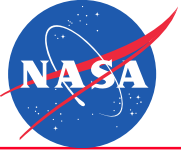
where x is the number of observed counts in n quarter frames.



- However, due to the modulated signal, the blocking compensation has to be done on a slot-by-slot basis (i.e., on the readout FPGA). Thus, the blocking compensation is implemented using a lookup table using the correction:

$$x_{corr}(x) = \text{floor} \left(n_{slot} \left(-\ln \left(1 - \frac{x}{n_{slot} - x(N_{wait} - 1)} \right) \right) \right),$$

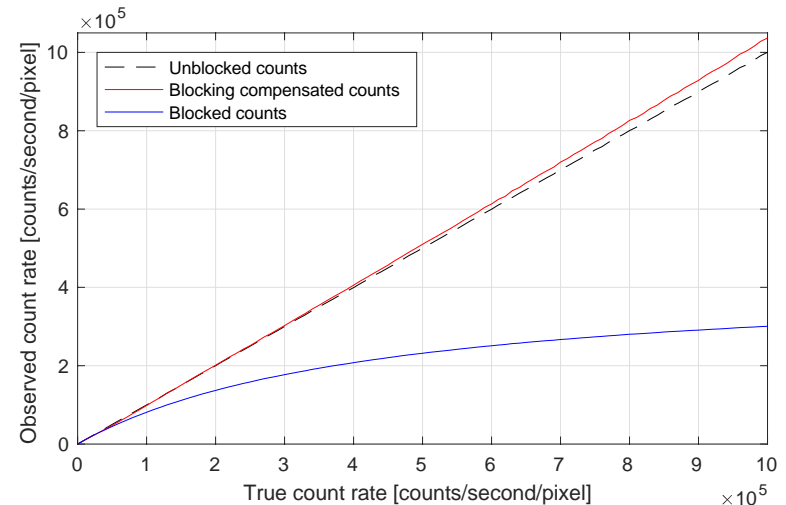
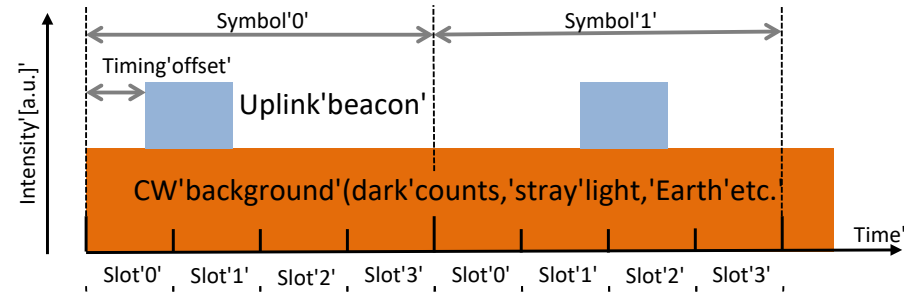
where x_{corr} is the corrected number of counts in one slot and n_{slot} is the number of quarter frames in a slot ($n_{slot} = 128$).



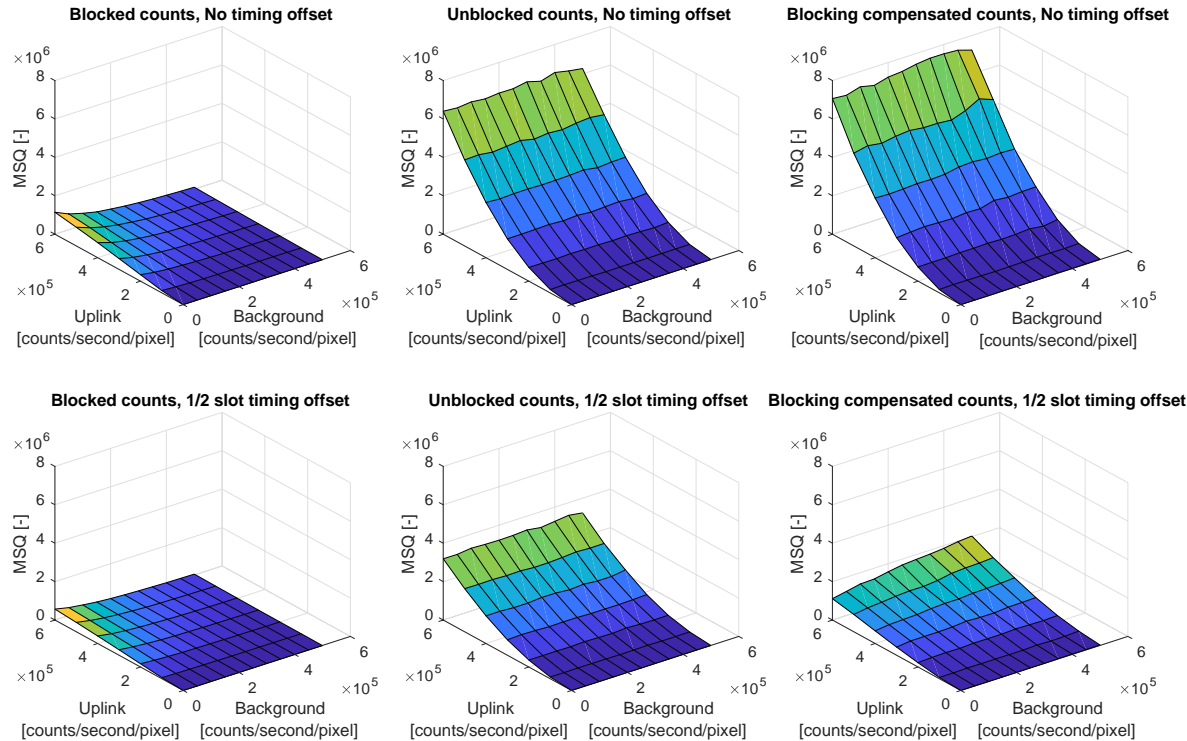
Blocking compensation cont.

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- The blocking compensation scheme assumes the count rate is constant over the integration time (one slot).
 - This assumption is reasonable in tracking mode where the timing offset is actively minimized
 - In acquisition mode the timing offset is unknown (1/2 slot worst case)
- Blocking compensation linearizes the detector response
 - This helps the underlying algorithms (e.g., the MSQ statistic) which were derived assuming a linear response
 - It does not “create” signal where there was none. I.e., the noise is also amplified.
- The limited integration time and restriction to integer numbers causes a small overestimation of the corrected counts.



Monte Carlo simulations the impact of blocking on the MSQ statistic under relevant uplink and background count rates



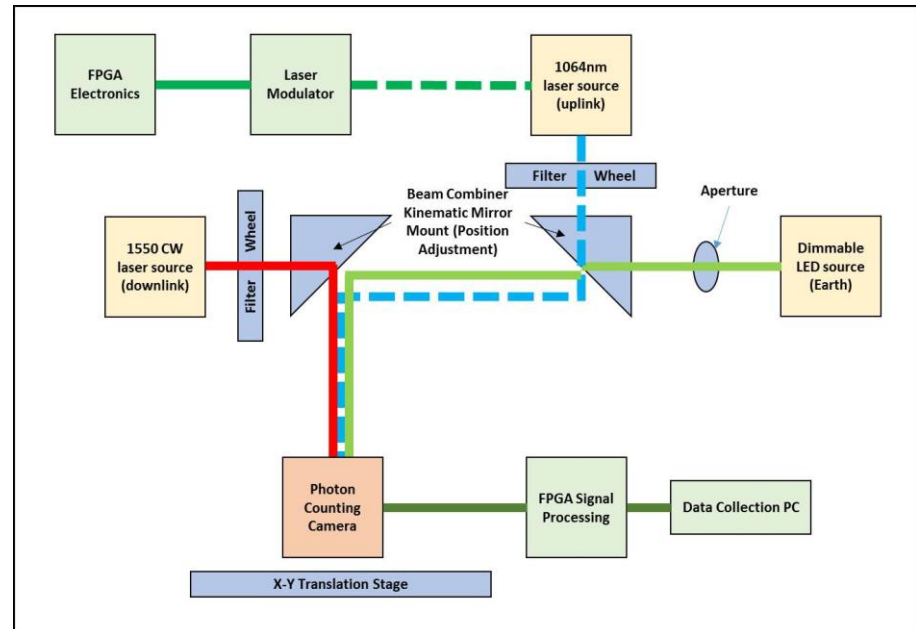
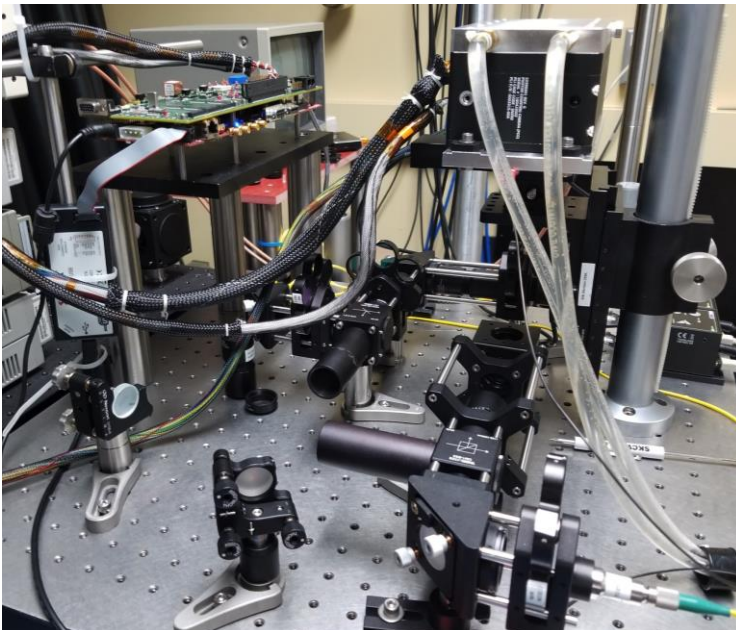
No timing offset: blocking compensation works well

Worst case timing offset: As expected blocking compensation is not ideal, but meets Acquisition mode requirements

- Contrary to previous assumptions, blocking compensation is important, even in the absence of background.

Centroiding performance was tested using two models:

- A high fidelity Monte Carlo code, simulating the detector array, ROIC, and readout FPGA in software
- A Photon Counting Camera test bed for detector and algorithm testing under realistic conditions
 - Including separate 1064nm uplink and 1550nm downlink sources, variable background models
 - PCC mounted on motorized precision XY translation stage for accurate (relative) movements.

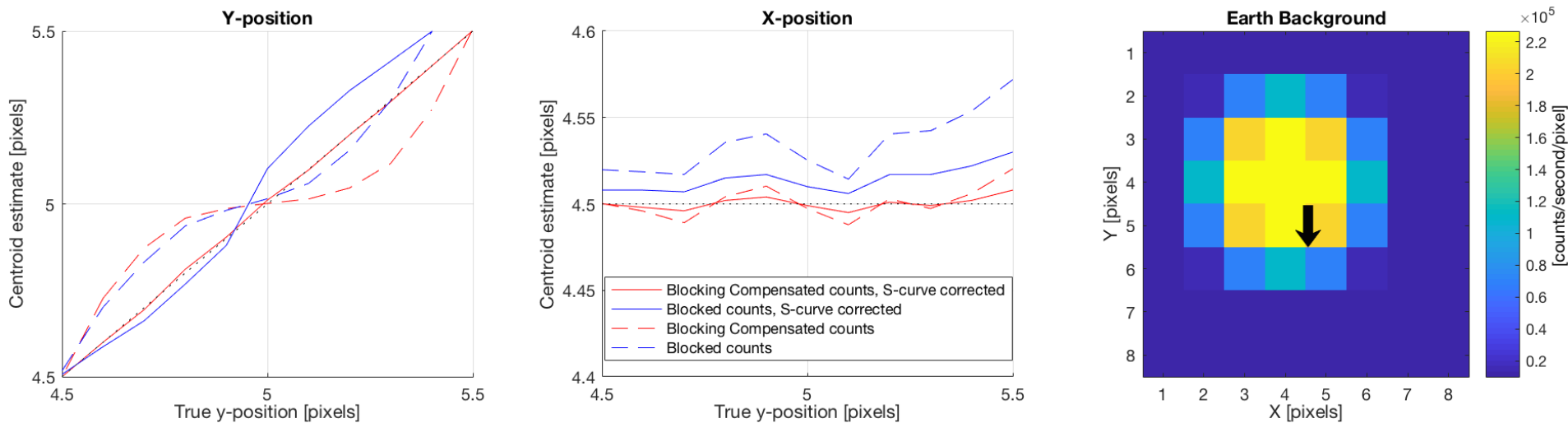




Performance testing cont.

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- A standardized test case was constructed and used as the baseline for testing.
- This test case represent the worst case conditions for the DSOC mission:
 - 2.7 AU Earth-Spacecraft distance
 - Fully illuminated Earth background ($0.0087 \text{ W}/(\text{cm}^2 \cdot \text{sr} \cdot \mu\text{m})$) with a 1 nm bandpass filter
 - Low beacon power at detector ($\sim 100 \text{ fW}$ average power)
 - 30% detector efficiency assuming 100% fill factor
 - 10000 counts/second/pixel Dark counts and Stray light
 - Centroid update rate 59.6 Hz (64 symbols integration time)
 - 5 quarter frames blocking time.



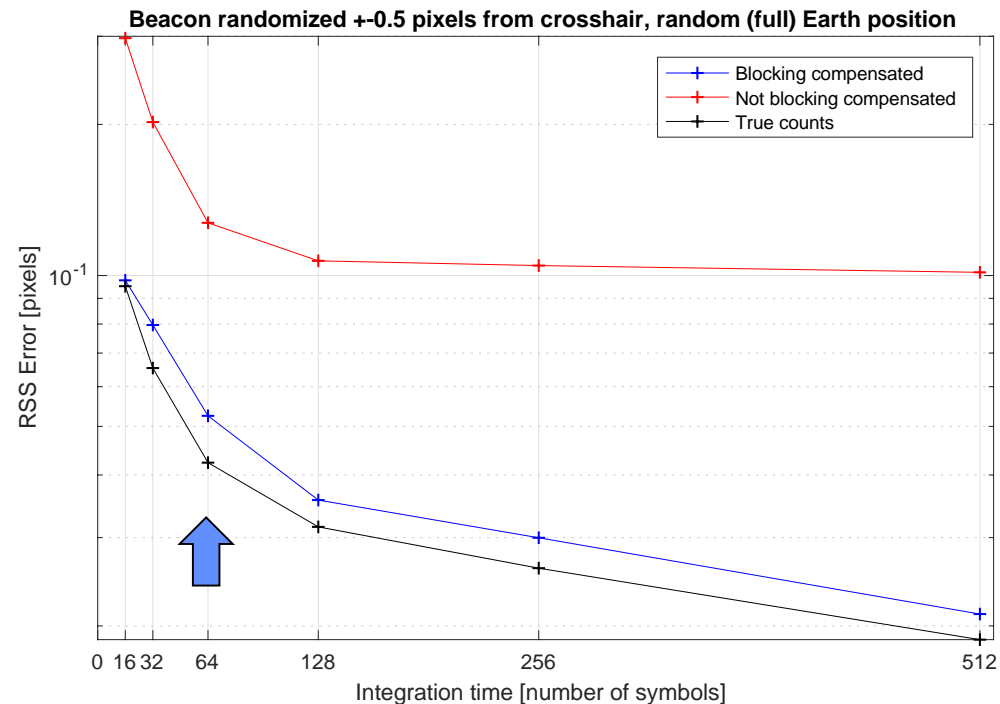
- Using the 2.7 AU test case the uplink beacon centroid is calculated for various Y-position on a full Earth background (indicated by the arrow), using a 3x3 subwindow.
- Using the blocking compensated counts, the Y-centroid estimate shows a symmetric S-curve (red dashed) which is easily corrected for (red solid).
- Without blocking compensation, the X- and Y-centroid estimates show significant asymmetry (blue dashed). This is due to the background dependence of the MSQ statistic resulting in a background dependent bias. This bias can not be corrected for (blue solid).
- The integration time here is 50x64 symbols (0.84s) to highlight bias instead of noise.



Monte Carlo Centroid tests cont.

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- The Monte Carlo code was used for a wide range of parametric studies
- Here the RSS error is shown as a function of the integration time
 - (64 symbols nominal)
- The results for the blocking compensated counts follow the true counts with a small offset:
 - The RSS error decreases with longer integration time.
- However, the RSS error for the non-blocking compensated counts level out at ~0.1 pixels RSS error due to the limit imposed by bias error.



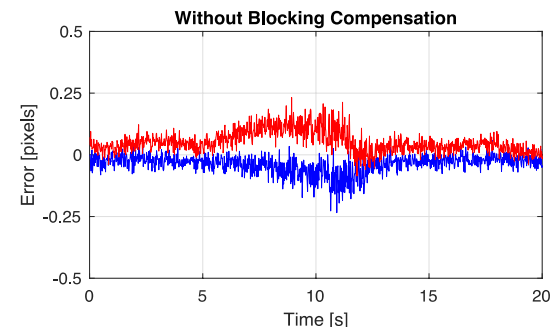
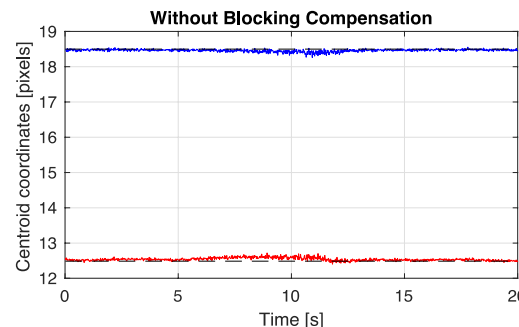
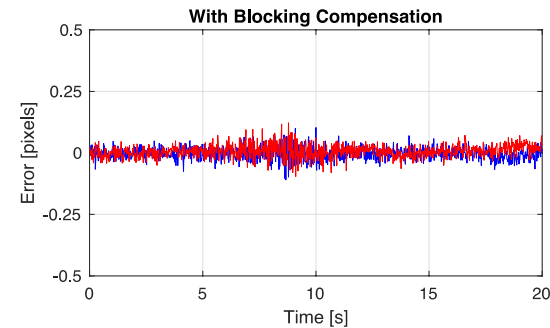
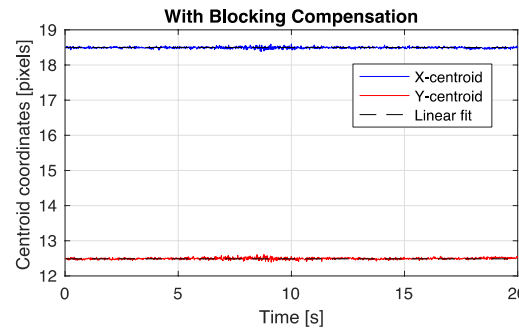
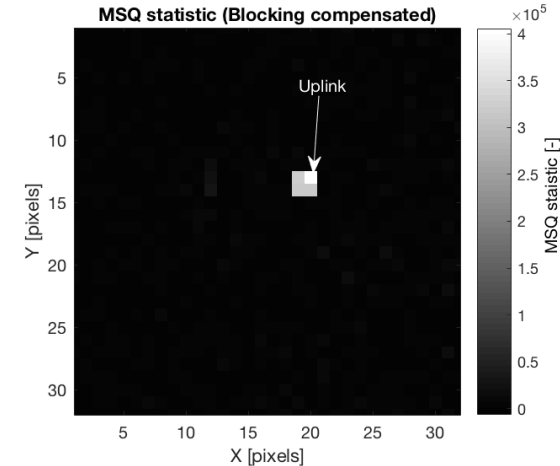
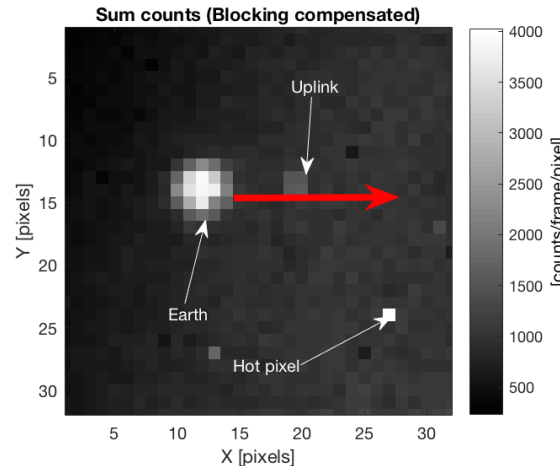
- 2.7 AU test case
- 2x2 centroiding subwindow
- S-curve correction applied
- No timing offset
- $RSS = \sqrt{RMS_X^2 + RMS_Y^2}$



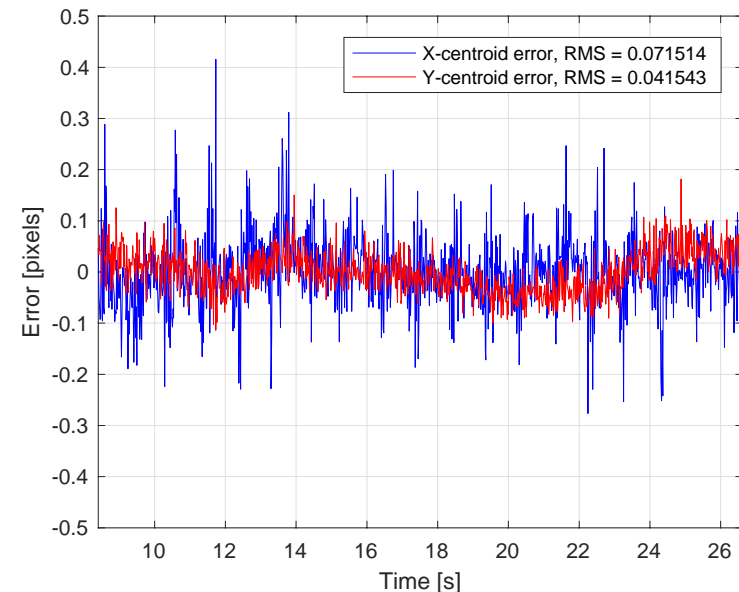
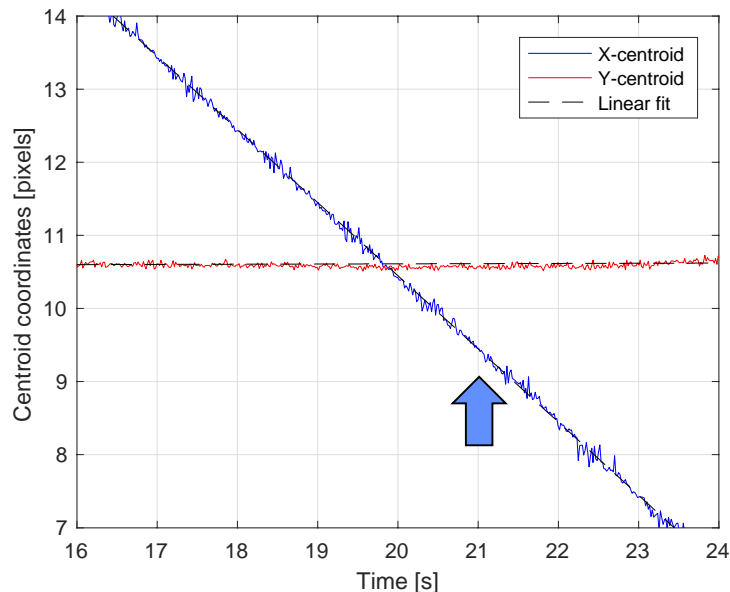
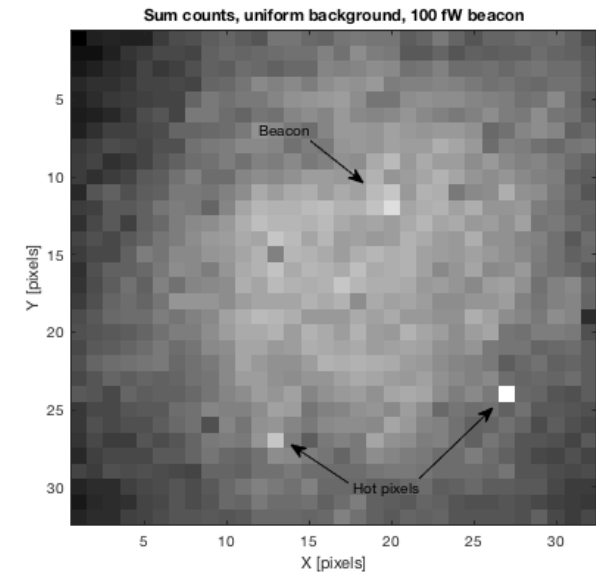
Laboratory Centroid test

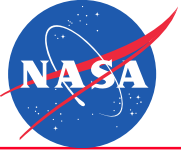
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- In the laboratory test setup it is easier to move the Earth over a stationary uplink spot
- Here we show the X- and Y-centroid estimates for two separate Earth sweeps, with, and without blocking compensation enabled.
- Signal and background levels are consistent with the 2.7AU test case.
- With blocking compensation enabled the centroid estimate jitter increases with the Earth present, but the mean position stays constant (no bias)
- Without blocking compensation, significant bias is apparent, as expected.



- Lacking an absolute beacon position reference, linear sweeps of the beacon spot over a uniform background at a constant rate gives a relative reference.
- A linear fit to the X and Y centroid estimates thus acts as the “true” position.
- X- and Y-centroid errors are 0.07 and 0.04 pixel RMS respectively over the entire sweep.
- In tracking mode we would choose to operate over a 4-pixel crosshair where jitter is minimized.





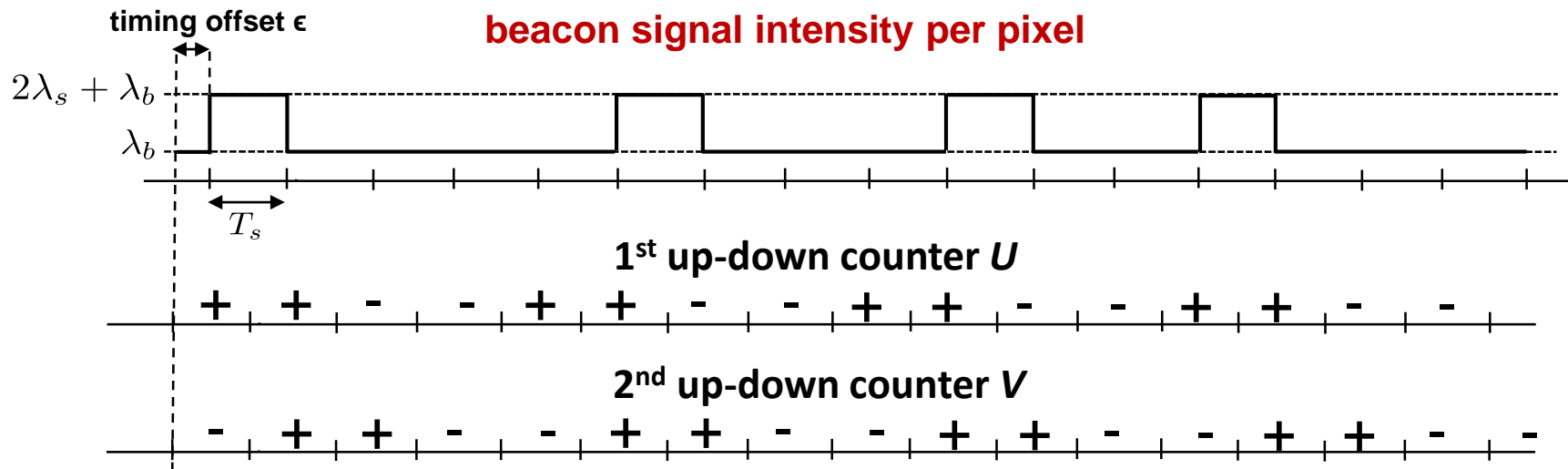
Summary

- **Developed a blocking compensation model and implemented blocking compensation in PCC readout electronics (FPGA)**
- **Using Monte Carlo simulations and a laboratory test bed we showed that:**
 - **Blocking breaks the background independence of the MSQ statistic, leading to significant centroiding errors (bias)**
 - **Blocking compensated counts closely follow the True counts, with a small noise penalty associated with the discarding of data**
 - **Using the blocking compensation, combined with S-curve correction, leads to centroid estimates that are linear with displacement and free of bias.**
 - **Uplink beacon centroid jitter is $<1 \mu\text{rad}$ even under worst case dynamic conditions**
 - **In Tracking mode, near 4 pixel crosshair, uplink beacon jitter is $<0.5 \mu\text{rad}$**
- **The blocking compensation algorithm degrades with timing offset**
 - **Only matters in Acquisition mode where centroiding accuracy is secondary.**



Backup

- Up-down counters alternately increment and decrement pixel counts at beacon frequency



- The up-down counter (UDC) outputs U and V have mean and variance

no dependence on background

$$E[U] = 4N(1 - \epsilon)\lambda_s T_s, \quad Var[U] = 4N(\lambda_b + \lambda_s)T_s$$

$$E[V] = 4N\epsilon\lambda_s T_s, \quad Var[V] = 4N(\lambda_b + \lambda_s)T_s$$

where N is the number of beacon cycles, ϵ is the slot timing offset, and T_s is the slot time.

- Modified square law statistic $W^* = U^2 + V^2 - 2S$, where S is up-counts, enables signal detection without phase synchronization, and has expected value that does not depend upon background.